

## Homework #1

Problem 1. (a) Clearly

$$\begin{aligned}\nabla \times (\mathcal{L}u) &= \nabla \times \left( -\hat{z} \times u + \cancel{\nabla q}^0 \right) \\ &= (\hat{z} \cdot \nabla)u\end{aligned}$$

by vector calculus identities. Likewise,

$$-\hat{z} \times (\nabla \times u) = (\hat{z} \cdot \nabla)u - \nabla u_z.$$

Since  $(\hat{z} \cdot \nabla)u$  is divergence-free, we see that  $q = u_z$  and

$$\begin{aligned}\mathcal{L}(\nabla \times u) &= -\hat{z} \times (\nabla \times u) + \nabla u_z \\ &= (\hat{z} \cdot \nabla)u \\ &= \nabla \times (\mathcal{L}u).\end{aligned}$$

(b) Since  $\mathbf{k} \times \hat{z}^\perp = \mathbf{k} \times \hat{z}$

$$\begin{aligned}i\mathbf{k} \times h_s(\mathbf{k}) &= i\mathbf{k} \times \hat{z} - s\mathbf{k} \times (\hat{\mathbf{k}} \times \hat{z}) \\ &= s|\mathbf{k}| \left[ is\hat{\mathbf{k}} \times \hat{z} - \hat{\mathbf{k}} \times (\hat{\mathbf{k}} \times \hat{z}) \right] \\ &= s|\mathbf{k}| \left[ \hat{z}^\perp + is\hat{\mathbf{k}} \times \hat{z} \right] \\ &= s|\mathbf{k}| h_s(\mathbf{k})\end{aligned}$$

$$\text{Since } \hat{\mathbf{z}} \times \hat{\mathbf{z}}^\perp = -(\hat{\mathbf{z}} \cdot \hat{\mathbf{k}}) \hat{\mathbf{z}} \times \hat{\mathbf{k}}$$

$$\begin{aligned} -\hat{\mathbf{z}} \times h_s(\mathbf{k}) &= (\hat{\mathbf{z}} \cdot \hat{\mathbf{k}}) \hat{\mathbf{z}} \times \hat{\mathbf{k}} - is \hat{\mathbf{z}} \times (\hat{\mathbf{k}} \times \hat{\mathbf{z}}) \\ &= -(\hat{\mathbf{z}} \cdot \hat{\mathbf{k}}) \hat{\mathbf{k}} \times \hat{\mathbf{z}} + is \hat{\mathbf{z}} \times (\hat{\mathbf{z}} \times \hat{\mathbf{k}}) \\ &= is (\hat{\mathbf{z}} \cdot \hat{\mathbf{k}}) \times [is \hat{\mathbf{k}} \times \hat{\mathbf{z}}] \\ &\quad + is [-\hat{\mathbf{k}} + (\hat{\mathbf{z}} \cdot \hat{\mathbf{k}}) \hat{\mathbf{z}}] \\ &= is (\hat{\mathbf{z}} \cdot \hat{\mathbf{k}}) [\hat{\mathbf{z}} + is \hat{\mathbf{k}} \times \hat{\mathbf{z}}] - is \hat{\mathbf{k}} \\ &= is (\hat{\mathbf{z}} \cdot \hat{\mathbf{k}}) [\hat{\mathbf{z}}^\perp + is \hat{\mathbf{k}} \times \hat{\mathbf{z}}] \\ &\quad - is [1 - (\hat{\mathbf{z}} \cdot \hat{\mathbf{k}})^2] \hat{\mathbf{k}} \\ &= is (\hat{\mathbf{z}} \cdot \hat{\mathbf{k}}) h_s(\mathbf{k}) - ik \hat{q}(\mathbf{k}) \end{aligned}$$

so that

$$\begin{aligned} \mathcal{L} h_s(\mathbf{k}) &= -\hat{\mathbf{z}} \times h_s(\mathbf{k}) + ik \hat{q}(\mathbf{k}) \\ &= is (\hat{\mathbf{z}} \cdot \hat{\mathbf{k}}) h_s(\mathbf{k}), \quad \checkmark \end{aligned}$$

Problem 2 (a) Simple direct calculations give

$$\begin{aligned} E_{\langle \ell \rangle}(t) &= \frac{1}{2} \langle \tau_{\ell}(u_i, u_i) \rangle \\ &= \frac{1}{2} \left[ \langle |u|^2 \rangle - \langle |\bar{u}_{\ell}|^2 \rangle \right] \\ &= \frac{1}{2} \sum_{\mathbf{k}} \left[ 1 - \left| (\pi)^2 \hat{G}_{\ell}^P(\mathbf{k}) \right|^2 \right] |\hat{u}(\mathbf{k}, t)|^2 \end{aligned}$$

where we used

$$\begin{aligned} \bar{u}_{\ell}(x) &= \int_{\mathbb{T}^2} d^2 r G_{\ell}^P(r) u(x+r) \\ &= (\pi)^2 (G_{\ell}^P * u)(r) \quad \left( \begin{array}{l} \text{assuming for} \\ \text{simplicity that} \\ G \text{ is even} \end{array} \right) \end{aligned}$$

and then applying the convolution theorem for Fourier transforms on  $\mathbb{T}^2$ . Furthermore, the Poisson summation formula with our definitions of Fourier transform in the course gives

$$(\pi)^2 \hat{G}_{\ell}^P(\mathbf{k}) = \hat{G}_{\ell}(\mathbf{k}) = \hat{G}(\ell \mathbf{k})$$

implying

$$E_{\langle \ell \rangle}(t) = \frac{1}{2} \sum_{\mathbf{k} \in \mathbb{Z}^2} \left[ 1 - |\hat{G}(\ell \mathbf{k})|^2 \right] |\hat{u}(\mathbf{k}, t)|^2$$

Note that for all  $k \in \mathbb{R}^2$

$$|\hat{G}(k)| = \left| \int_{\mathbb{R}^2} d^2r G(r) e^{-ik \cdot r} \right|$$

$$\leq \int_{\mathbb{R}^2} d^2r G(r) = 1$$

Hence,

$$\sup_{|k| \geq 1} \frac{1 - |\hat{G}(k)|^2}{|k|^2} \leq 1$$

whereas using  $|\hat{G}(k)|^2 = 1 - O(|k|^{-2})$ , i.e.  $1 - |\hat{G}(k)|^2 < C|k|^{-2}$ ,

$$\sup_{|k| \leq 1} \frac{1 - |\hat{G}(k)|^2}{|k|^2} \leq C.$$

It follows that

$$E_{<\ell}(t) = \frac{\ell^2}{2} \sum_{k \in \mathbb{Z}^2} \frac{1 - |\hat{G}(\ell k)|^2}{|\ell k|^2} \cdot |\bar{\omega}(k, t)|^2$$

$$\leq \ell^2 \left[ \sup_{k \in \mathbb{R}^2} \frac{1 - |\hat{G}(k)|^2}{|k|^2} \right] \cdot \sum_{k \in \mathbb{Z}^2} \frac{1}{2} |\bar{\omega}(k, t)|^2$$

$$= \ell^2 \max\{1, C\} \Omega(t) \quad \checkmark$$

(b) A similar calculation as in (a) shows that

$$\begin{aligned}
 \Omega_{>\ell}(t) &= \frac{1}{2} \langle |\bar{w}_\ell|^2 \rangle \\
 &= \frac{1}{2} \sum_{\mathbf{k}} \left| (\pi\ell)^2 \hat{G}_\ell^D(\mathbf{k}) \right|^2 |\hat{w}(\mathbf{k}, t)|^2 \\
 &= \frac{1}{2} \sum_{\mathbf{k}} \left| \hat{G}(\ell\mathbf{k}) \right|^2 \cdot |\mathbf{k}|^2 |\hat{u}(\mathbf{k}, t)|^2 \\
 &= \frac{1}{2\ell^2} \sum_{\mathbf{k}} |\ell\mathbf{k} \cdot \hat{G}(\ell\mathbf{k})|^2 |\hat{u}(\mathbf{k}, t)|^2
 \end{aligned}$$

Note that for all  $\mathbf{k} \in \mathbb{R}^2$

$$i\mathbf{k} \hat{G}(\mathbf{k}) = \int_{\mathbb{R}^2} d^2x \nabla G(\mathbf{x}) e^{-i\mathbf{k} \cdot \mathbf{x}}$$

$\Rightarrow$

$$|\mathbf{k} \hat{G}(\mathbf{k})| \leq \int_{\mathbb{R}^2} d^2x |\nabla G(\mathbf{x})| = \|\nabla G\|_{L^1} < +\infty$$

and thus

$$\begin{aligned}
 \Omega_{>\ell}(t) &\leq \frac{\|\nabla G\|_{L^1}^2}{\ell^2} \sum_{\mathbf{k}} \frac{1}{2} |\hat{u}(\mathbf{k}, t)|^2 \\
 &= \frac{C}{\ell^2} E(t). \quad \checkmark
 \end{aligned}$$

Problem 3 (a) The total helicity in wavenumbers  $< k$  can be bounded as

$$\begin{aligned}
 |H_{<k}(t)| &= \left| \int_0^k dk H(k,t) \right| \\
 &\leq \int_0^k dk |H(k,t)| \\
 &\leq \int_0^k dk 2k E(k,t) \\
 &\leq 2k \int_0^k dk E(k,t) \\
 &\leq 2k \int_0^\infty dk E(k,t) = 2k \cdot E(t)
 \end{aligned}$$

However, this is a bit misleading, because  $H(k,t)$  is a signed quantity and thus one could have subintervals  $[k_1, k_2]$

$$\left| \int_{k_1}^{k_2} dk H(k,t) \right| \gg 2k \cdot E(t)$$

as long as there is another interval  $[k_3, k_4] \subseteq [0, k] \setminus [k_1, k_2]$  with equal helicity of the opposite sign.

For energy there is no upper bound at all on the amount to reach wavenumbers  $> K$ , because there is only a lower bound

$$\begin{aligned}
 E_{>K}(t) &= \int_K^{\infty} dk E(k,t) \\
 &\geq \int_K^{\infty} dk \frac{|H(k,t)|}{2k} \\
 &\geq \frac{1}{2K} \int_K^{\infty} dk |H(k,t)| \\
 &\geq \frac{1}{2K} \left| \int_K^{\infty} dk H(k,t) \right|.
 \end{aligned}$$

Furthermore,  $\int_K^{\infty} dk H(k,t)$  is not conserved in time, even if the lower bound were an upper bound.

(b) For 3D NS projected onto the  $\pm$ -helicity modes, the situation is quite different. For a mode  $h_{\pm}(k)$

$$ik \times h_{\pm}(k) = k h_{\pm}(k)$$

and thus

$$i\mathbf{k} \times \mathbf{h}_+(\mathbf{k}) \cdot \mathbf{h}_+^*(\mathbf{k}) = k |\mathbf{h}_+(\mathbf{k})|^2.$$

Summing over all such modes with fixed wavenumber, one obtains now that

$$H(\mathbf{k}, t) \geq 0 \quad \text{and} \quad H(\mathbf{k}, t) = 2k E(\mathbf{k}, t).$$

The situation is therefore now exactly analogous to 2D! One can see first of all that

$$H_{<K}(t) \leq 2K \cdot E(t)$$

and also that

$$\begin{aligned} E_{>K}(t) &= \int_K^\infty dk \frac{H(\mathbf{k}, t)}{2k} \\ &\leq \frac{1}{2K} \int_0^\infty dk H(\mathbf{k}, t) = \frac{1}{2K} H(t). \end{aligned}$$

Furthermore,  $E(t) = E_0$  and  $H(\mathbf{k}) = H_0$ , because both quantities are conserved in detail for individual helical modes and thus for the dynamics including all the triadic interactions with positive helicity modes.

Problem 4, (a) Since  $e^{\beta\psi(r)} = \frac{1}{(1-Ar^2)^2}$ , one finds

$$Z = \int_0^1 2\pi r dr e^{\beta\psi(r)}$$

$$= \pi \int_0^1 \frac{2r dr}{(1-Ar^2)^2} \quad \text{with } u=r^2$$

$$= \pi \int_0^1 \frac{du}{(1-Au)^2} = \frac{\pi}{A} \left. \frac{1}{1-Au} \right|_{u=0}^{u=1} = \frac{\pi}{1-A}$$

Furthermore,

$$u_\theta = \frac{\partial \psi}{\partial r} = -\frac{2}{\beta} \frac{1}{1-Ar^2} (-2Ar) = \frac{4A}{\beta} \frac{r}{1-Ar^2} \checkmark$$

and thus

$$ru_\theta = r \frac{\partial \psi}{\partial r} = \frac{4}{\beta} \frac{Ar^2}{1-Ar^2} = \frac{4}{\beta} \left( -1 + \frac{1}{1-Ar^2} \right)$$

So that

$$\omega = \frac{1}{r} \frac{\partial}{\partial r} (ru_\theta) = \frac{1}{r} \cdot \frac{4}{\beta} \cdot \frac{-1}{(1-Ar^2)^2} \cdot (-2Ar)$$

$$= \frac{8A}{\beta} \frac{1}{(1-Ar^2)^2} = \frac{1-A}{\pi} \frac{1}{(1-Ar^2)^2} \checkmark$$

using  $1-A = \frac{8\pi}{\beta} A$ . Finally,  $\omega = \Delta\psi = \frac{1}{Z} e^{\beta\psi}$ .

(b) Clearly

$$w(x) \geq 0$$

inside the disk  $D$ . Also,

$$\int_D w(x) d^2x = \frac{\int_D e^{\beta \varphi} d^2x}{Z} = \frac{Z}{Z} = 1.$$

Furthermore, it is advantageous to write  $w$  in the form

$$w(r) = \frac{8}{|\beta|} \frac{r_1^2}{(r^2 + r_1^2)^2}.$$

We then see that for any  $0 < \rho < 1$  and  $\beta < -4\pi$

$$w(r) < \frac{2}{\pi} \frac{r_1^2}{\rho^2} \quad \text{for } r > \rho$$

$$\longrightarrow 0$$

as  $\beta \rightarrow -8\pi$  and  $r_1 \rightarrow 0$ . Thus,  $w$ , labelled by  $\beta$

forms a standard approximating sequence of delta

functions as  $\beta \rightarrow -8\pi$ . In particular, for any smooth

test function

$$\lim_{\beta \rightarrow -8\pi} \int_D \varphi(x) w_\beta(x) d^2x = \varphi(0).$$

The energy is calculated as

$$\begin{aligned} E &= \frac{1}{2} \int_D |u|^2 d^2x && \text{since } u_r = 0 \\ &= \frac{1}{2} \int_0^1 2\pi r dr \left( \frac{4A}{\beta} \frac{r}{1-Ar^2} \right)^2 \\ &= \frac{8\pi A}{\beta^2} \int_0^1 du \frac{Au}{(1-Au)^2} && \text{with } u=r^2 \\ &= \frac{8\pi A}{\beta^2} \int_0^1 du \left[ \frac{-1}{1-Au} + \frac{1}{(1-Au)^2} \right] \\ &= \frac{8\pi}{\beta^2} \left[ \ln(1-A) - \frac{A}{1-A} \right] \end{aligned}$$

$$\sim \frac{1}{8\pi} \ln|A| \quad \text{as } \beta \rightarrow -8\pi, A \rightarrow -\infty$$

$$\sim -\frac{1}{4\pi} \ln r_1 \quad \text{using } r_1 = \frac{1}{\sqrt{A}}$$

so that

$$r_1 \sim e^{-4\pi E}$$

at high energies.

(c) We see that

$$u_\theta = \frac{\partial \psi}{\partial r} = \begin{cases} \frac{1}{2} \omega_0 r & r < r_1 \\ \frac{1}{2} \omega_0 r_1^2 \cdot \frac{1}{r} & r > r_1 \end{cases}$$

and

$$\omega = \frac{1}{r} \frac{\partial}{\partial r} (r u_\theta) = \begin{cases} \omega_0 & r < r_1 \\ 0 & r > r_1 \end{cases}$$

which is indeed a circular vortex patch. The normalization

$\int_D \omega d^2x = 1$  requires that  $\omega_0 = 1/\pi r_1^2$ , so that

$$u_\theta = \begin{cases} \frac{1}{2\pi} \frac{r}{r_1^2} & r < r_1 \\ \frac{1}{2\pi} \frac{1}{r} & r > r_1 \end{cases}$$

The energy has two parts

$$E^< = \frac{1}{2} \int_0^{r_1} \frac{1}{(2\pi r_1^2)^2} r^2 \cdot 2\pi r dr = \frac{1}{4\pi r_1^4} \int_0^{r_1} r^3 dr = \frac{1}{16\pi}$$

and

$$E^> = \frac{1}{2} \frac{1}{(2\pi)^2} \int_{r_1}^{\infty} \frac{1}{r^2} \cdot 2\pi r dr = \frac{1}{4\pi} \int_{r_1}^{\infty} \frac{dr}{r} = -\frac{1}{4\pi} \ln r_1$$

The second dominates for small  $r_1$ , so that  $E \sim -\frac{1}{4\pi} \ln r_1$ .

Problem 5. (a) Two wavenumbers  $p$  and  $q$  that interact to transfer enstrophy or energy to wavenumber  $k$  must satisfy

$$k = p + q$$

and thus

$$|k| \leq |p| + |q| < 2K$$

if  $|p|, |q| < K$ .

(b) The quantity  $\eta = \nu \langle |\nabla w|^2 \rangle$  has dimension

$$\begin{aligned} [\eta] &= \frac{(\text{length})^2}{\text{time}} \cdot \left( \frac{1}{\text{length} \cdot \text{time}} \right)^2 \\ &= \frac{1}{(\text{time})^3} \end{aligned}$$

Hence, the turnover rate is determined dimensionally to be

$$T_k \approx \eta^{-1/3}$$

independent of wavenumber  $k$ . This is quite distinct from the 3D energy cascade which is "accelerated," so that turnover times decrease as a power of  $\frac{1}{k}$  for  $k \rightarrow \infty$ .

(c) The number of steps required is  $\log_2 \left( \frac{k_2}{k_1} \right) = \frac{\ln(k_2/k_1)}{\ln 2}$  and the time required is

$$T \approx \eta^{-1/3} \ln \left( \frac{k_2}{k_1} \right).$$

Conversely, the wavenumber reached in time  $t$  will be

$$\begin{aligned}k_2(t) &= 2^{t/\tau} \cdot k_1 \\ &= k_1 \exp\left(\frac{\ln 2}{\tau} t\right) \\ &\approx k_1 \exp(\eta^{1/3} t).\end{aligned}$$

(d) In the  $k^{-3}$  range

$$\frac{1}{2} w_{rms}^2 = \int_{k_1}^{k_2} \eta^{2/3} \frac{dk}{k} \approx \eta^{2/3} \ln\left(\frac{k_2}{k_1}\right)$$

so that

$$w_{rms} \approx \eta^{1/3} \left[ \ln\left(\frac{k_2}{k_1}\right) \right]^{1/2}$$

(e) Using (c) and (d),

$$T \approx \frac{1}{w_{rms}} \left[ \ln\left(\frac{k_2}{k_1}\right) \right]^{3/2}$$

For the synoptic scales of Earth's atmosphere

$$\begin{aligned}T &\approx 10^4 \text{ sec} (\ln 4)^{3/2} \approx 1.63 \times 10^4 \text{ sec} \\ &\approx 4.53 \text{ hr}\end{aligned}$$